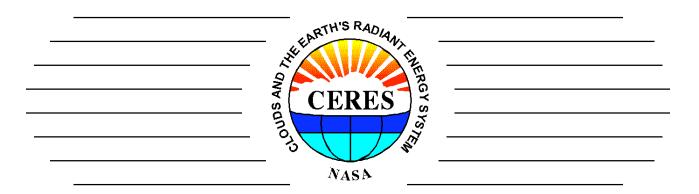
Terra/Aqua Instrument and ERBE-Like Calibration Status Report



Kory J. Priestley Susan Thomas, Denise Cooper Phil Hess, Grant Matthews, Peter Szewczyk, Dale Walikainen, Robert Wilson

CERES Science Team Meeting

Hampton, VA November 1-3, 2005





Agenda

CERES Project Status

Engineering/Housekeeping Status

FM-4 SW Channel Anomaly - Aqua Spacecraft

Cal/Val Report

Edition2_Rev1 Review

Preparations for Edition3 BDS & ERBE-like

- Spectral Darkening of CERES SW channels





CERES Terra/Aqua Health & Status

With the exception of the SW channel on the CERES/Aqua FM-4 Instrument, the CERES Terra/Aqua instruments are functioning nominally...

All temperatures and voltages remain within limits

No discernable or alarming trends.

Spacecraft	Instruments	Launch	Science Initiation	Collected Data (Months)
TRMM	PFM	11/97	1/98	9
Terra	FM1, FM2	12/99	3/00	68 +
Aqua	FM3, FM4	5/02	6/02	41 +
?	FM5	?	?	?





CERES/Aqua FM-4 SW Anomaly History & Status

History-

- On March 30, 2005 at 18:42 GMT the CERES/Aqua FM-4 instrument stopped collecting valid Shortwave channel radiometric measurements.
- Anomaly was characterized by an immediate railing 'high' of the SW channel data stream, which subsequently drove the SW channel bridge balance electronics into a reset condition in an attempt to bring the SW channel output on-scale.
- The Total and Atmospheric Window radiometric channels continued to function nominally.

Current Status-

- All attempts to recover the SW channel measurements have been unsuccessful.
- Modeling has indicated the problem is most likely the result of a failed resistor (possibly a detector) in the SW channel Bridge Balance Circuitry.
- •Preparing for upcoming Calipso launch.





CERES/Aqua FM-4 SW Anomaly Science Impacts

Near Term

No impact assuming the CERES/Aqua FM-3 instrument is collecting cross-track data.

Long Term

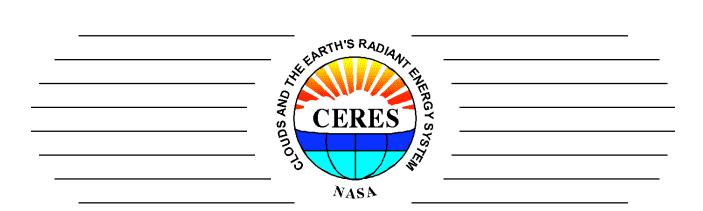
If anomaly cannot be resolved, longer term science issues include (in order of priority)

- significantly increased risk of a gap from Terra/Aqua climate record to the beginning of NPOESS using FM-5 in 2011/2012.
- the second Aqua CERES instrument for the A-train was to scan along-track for multiangle views over the lidar/radar track to examine effects like 3-D radiative transfer
- the second Aqua CERES instrument is also used to perform intercalibrations with GERB: but these can be accomplished by the second Terra CERES as long as two CERES instruments are active on Terra
- the second Aqua CERES instrument is also used during cloud/aerosol field experiments to provide multi-angle data that tracks a pre-selected surface site (e.g. ARM site).





Radiometric Performance Summary







Operational Coverage Request



http://asd-www.larc.nasa.gov/Instrument/operations.html





CERES BDS and ERBE-Like Product Status

Spacecraft	Product	Version	Available	Months Processed
TRMM	BDS	Edition1	Yes	1/98 - 8/98 , 3/00
	ERBE-Like	Edition1	Yes	1/98 - 8/98 , 3/00
		Edition2	Yes	1/98 - 8/98 , 3/00
Terra	BDS	Edition1	Yes	2/00 - present
		Edition2	Yes	2/00 - 6/05
	ERBE-like	Edition1	Yes	2/00 - present
		Edition2	Yes	2/00 - 6/05
Aqua	BDS	Edition1	Yes	6/02 - present
		Edition2	Yes	6/02 - 3/05
	ERBE-like	Edition1	Yes	6/02 - present
		Edition2	Yes	6/02 - 3/05

Note: Red text indicates months are in final validation prior to public release.





BDS and ERBE-Like Release Strategy

- Edition1 Static Algorithms and coefficients baseline product used in cal/val protocol
- Edition2 Utilizes temporally varying coefficients to correct for traceable radiometric drift. All spectral changes are broadband and 'gray'.
- Edition3 <u>Release date Fall 2006</u>. Will incorporate temporally varying spectral artifacts in the SW measurements. A complete re-analysis of Ground Calibration with additional component characterization measurements.

Revisions - Advance capabilities to the users prior to the release of the next Edition.

Edition2 products lag Edition1 by a minimum of 4 months, revisions may lag Edition releases.





Instrument Artifact Removal Strategy

Remote sensing instruments generally exhibit time varying artifacts in their data products. For CERES these artifacts stem from either of 2 physical entities.....

- Radiometric Gain Change
 - Wavelength independent change in sensor responsivity
 - Corrections implemented in Count Conversion algorithm (SS1)
- Spectral Response Change
 - Wavelength dependent change in sensor optics
 - Corrections implemented in Spectral Unfiltering algorithms (SS2)

The Edition2 production strategy assumes that spectral response changes are 'Gray' over broad spectral regions. In other words, we assume uniform spectral response changes in these broad regions.

Radiometric	Spectra	Spectral Region			
Channel	SW	LW			
Total	<3.0 um	>3.0 um			
sw	<5.0 um	-			
WN	-	8-12 um			





Notification of Revision



CERES BDS (BiDirectional Scan) Terra Edition2 Data Quality Summary

Investigation: **CERES**

Data Product: **BiD**irectional **S**can [BDS]
Data Set: **Terra (Instruments: FM1, FM2)**

Data Set Version: Edition2

The purpose of this document is to inform users of the accuracy of this data product as determined by the CERES Team. This document briefly summarizes key validation results, provides cautions where users might easily misinterpret the data, provides links to further information about the data product, algorithms, and accuracy, gives information about planned data improvements. This document also automates registration in order to keep users informed of new validation results, cautions, or improved data sets as they become available.

This document is a high-level summary and represents the minimum information needed by scientific users of this data product. It is strongly suggested that authors, researchers, and reviewers of research papers re-check this document for the latest status before publication of any scientific papers using this data product.

Table of Contents

- •Nature of the BDS Product
- •Updates to Current Edition
- •User Applied Revisions
- •Validation and Quality Assurance
- •Current Estimated Uncertainty of Data
- •Cautions When Using Data
- Expected Reprocesings
- •References
- •Web links to Relevant information
- •Referencing Data in Journal Articles
- •Giving Data to Other Users





BDS & ERBE-Like Edition2 Summary

•Cal/Val Protocol demonstrates radiometric stability of the Ed2 data products through 11/2004 of....

	Те	rra	Aqua		
	FM1	FM2	FM3	FM4	
LWday	.125	.125	<.1	<.1	
LWnight	<.1	<.1	<.1	<.1	
sw	.2	.3	~.1	~.1	
WN	<.1	<.1	.125	.125	

Note: Values apply to all-sky global averages

Units are in %/yr





BDS & ERBE-Like Edition2 Summary

•Cal/Val Protocol demonstrates radiometric stability of the Ed2 data products through 11/2004 of....

	Те	rra	Aqua		
	FM1	FM2	FM3	FM4	
LWday	.125	.125	<.1	<.1	
LWnight	<.1	<.1	<.1	<.1	
sw	.2	.3	~.1	~.1	
WN	<.1	<.1	.125	.125	

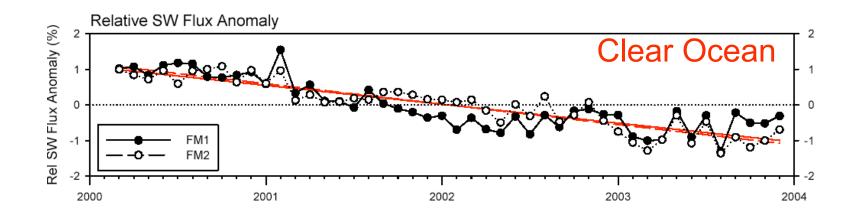
Note: Values apply to all-sky global averages

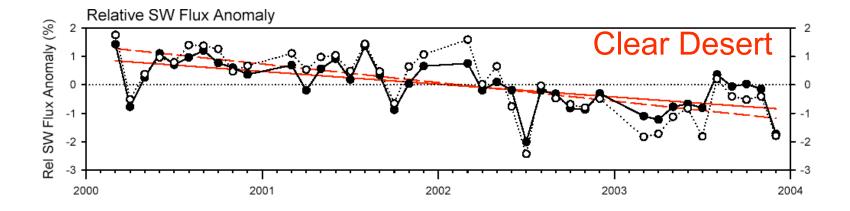
Units are in %/yr





CERES SSF Ed2B SW TOA Flux Anomaly









Direct Comparison of Nadir Radiance Measurements

Two CERES instruments on a common platform allows for a unique validation opportunity.....

Direct Comparison of simultaneous Nadir measurements

Each CERES/Terra instrument views nadir every 3.3 seconds

Thus, we obtain nearly simultaneous measurements of the same geolocation ($\Delta t < 3.3$ seconds)....

Spatial, angular, and temporal sampling issues are virtually eliminated.

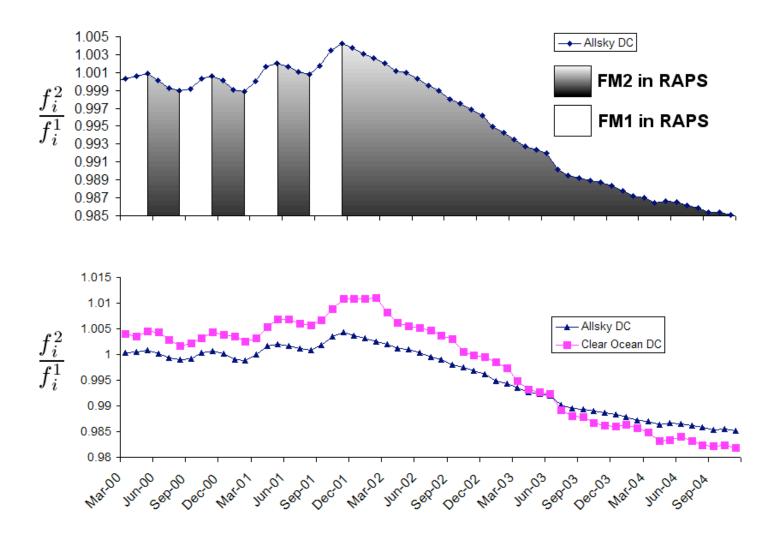
26,000 co-located (but not independent) measurements in a given day, provides a very rigorous statistical tool.

Results can be discretized by scene type to enhance the analysis.



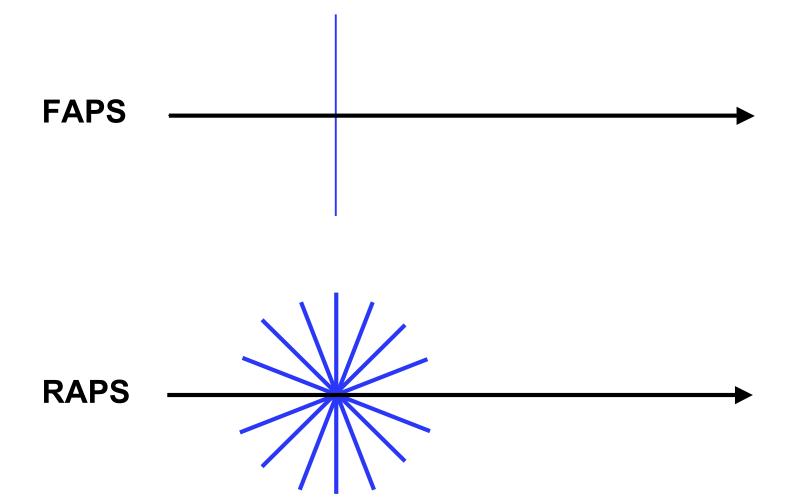


Edition2 BDS SW Nadir Direct Comparison



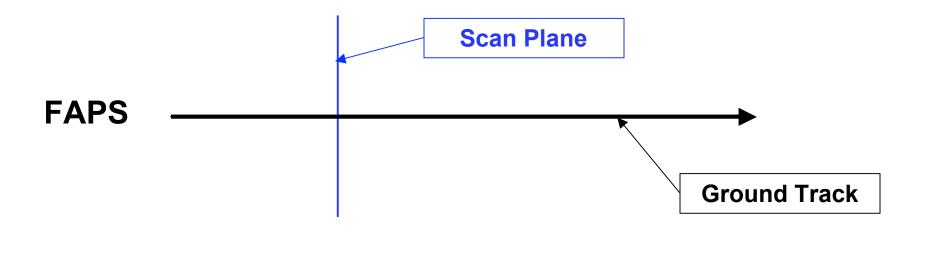


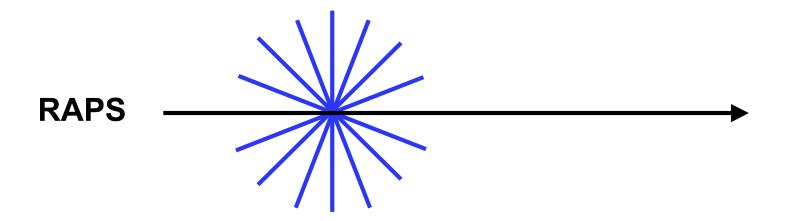






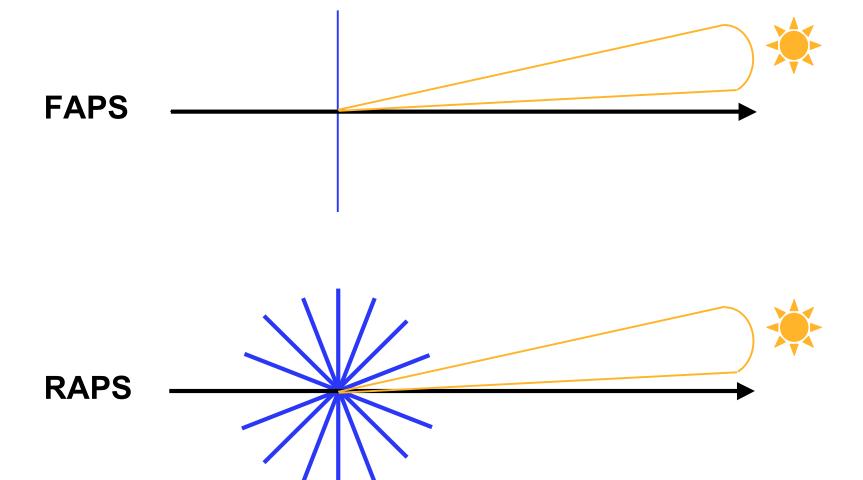






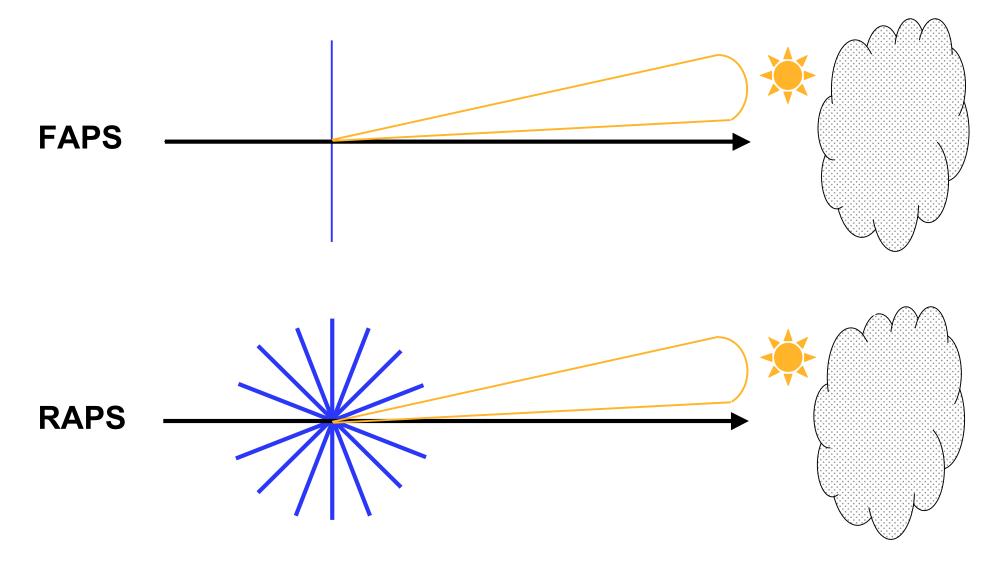
















Operations to Characterize Spectral Darkening

Month (2005)	FM1 Azimuth Gimbal	FM1 Solar Calibrations	FM2 Azimuth Gimbal	FM2 Solar Calibrations	FM2 Solar Exposure	FM2 RAM Direction	Darkening Rate Impact
February	Xtrack	Yes	Xtrack, Atrack	Yes	Nominal	Nominal (lunar scans)	Nominal
March	Xtrack	Yes	Xtrack	Yes (Daily)	Accelerated	Accelerated (solcal)	No Change
April	Xtrack	Yes	Rotating	Yes	Reduced (note1)	Nominal	No Change
Мау	Xtrack	Yes	Rotating, Atrack	Yes	Eliminated (note2)	Reduced (note2)	
June	Xtrack (note3)	Yes	Xtrack (note4)	No	Eliminated	Eliminated	
July	Xtrack (note3,5)	Yes	Stowed	No	-	-	
August	Xtrack	Yes	Stowed	No	-	-	
September	Xtrack	Yes	Xtrack (note6)	No	Eliminated	Eliminated	
October	Xtrack	No	Xtrack	No	Eliminated	Eliminated	
November	Xtrack	No	Xtrack	No	Eliminated	Eliminated	

Notes:

- 1. Short-Earth scan profile turn-around reduced 8 degrees on March 31, 2005 to both FM1 and FM2.
- 2. Expanded solar-avoidance region criteria May 13, 2005.
- 3. Greenland Summer Solstice Terra-Aqua Inter-Calibration, June 6- July 6, 2005
- 4. GERB Operations, June 9-30, 2005.
- 5. Nighttime internal calibrations, July 1-7, 2005.
- 6. Valencia Over-Flights, September 12-17, 2005.

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June	Xtrack (note3)	Yes	Xtrack (note4)	No	Eliminated	Eliminated	
July	Xtrack (note3,5)	Yes	Stowed	No	-	-	
August	Xtrack	Yes	Stowed	No	-	-	
September	Xtrack	Yes	Xtrack (note6)	No	Eliminated	Eliminated	
October	Xtrack	No	Xtrack	No	Eliminated	Eliminated	
November	Xtrack	No	Xtrack	No	Eliminated	Eliminated	

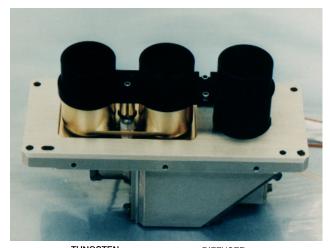
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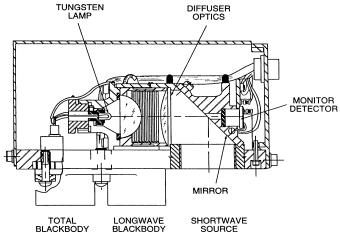
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CERES Onboard Calibration Sources

Internal Calibration Module (ICM)

- ICM Provides 3 unique radiance levels for the SW and LW sources
- Blackbodies for the Total and Window channels
- Temperature knowledge obtained via Platinum Resistance Thermometers
- Quartz-halogen tungsten lamp for the Shortwave channel (2100, 1900, 1700 K spectrums)
- SiPd independently monitors lamp output
- Acceptance testing for lamps includes burn-in period of ~80 hours to screen for stability
- Narrow and broadband radiant intensity measurements are performed prior to acceptance
- Design specification is +-0.5% stability over 5-year mission

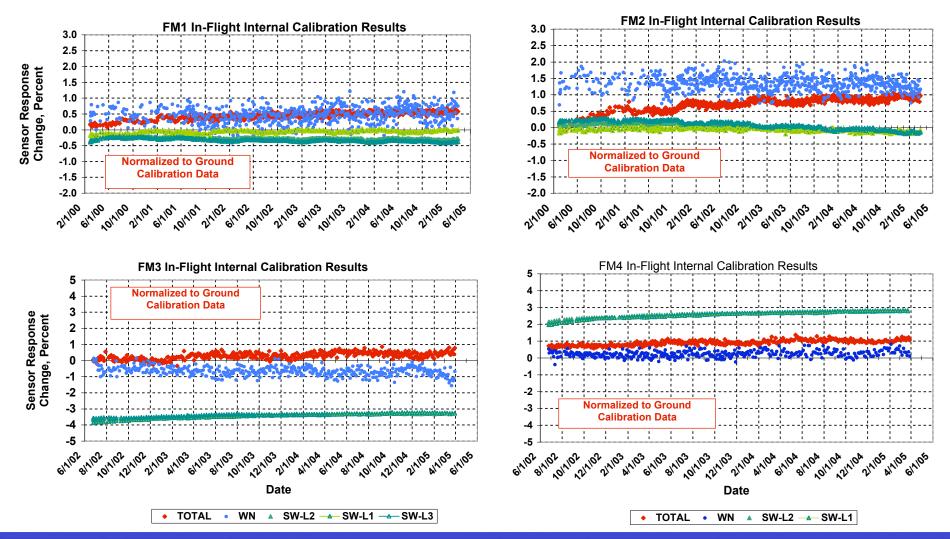








CERES Internal Calibration Results

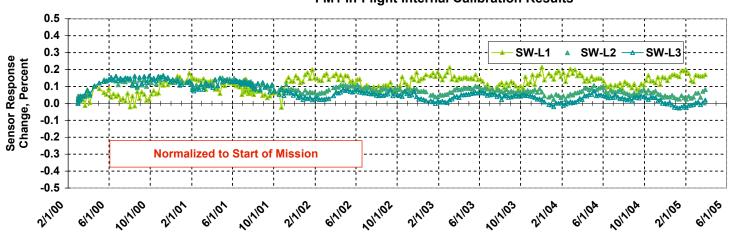




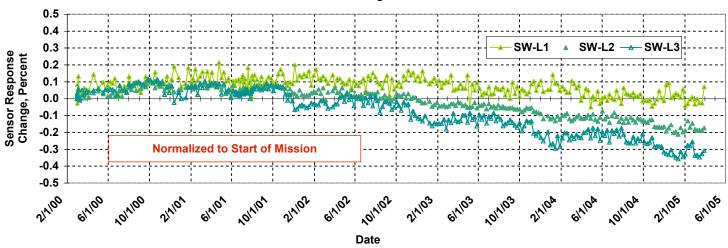


Terra SW Internal Calibration Results





FM2 In-Flight Internal Calibration Results



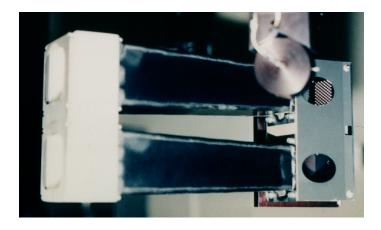


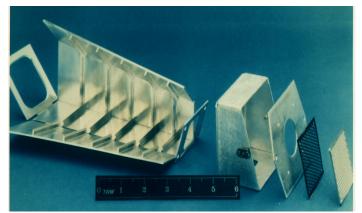


CERES Solar Calibration

Mirror Attenuator Mosaic (MAM)

- Solar Diffuser plate attenuates direct solar view
- MAM is a Nickel substrate with Aluminum coated spherical cavities or divots
- Provides a Relative calibration of the Shortwave channel and the SW portion of the Total channel
- Solar Cal results to date are suspect due to large initial drift in MAM surface reflectances...

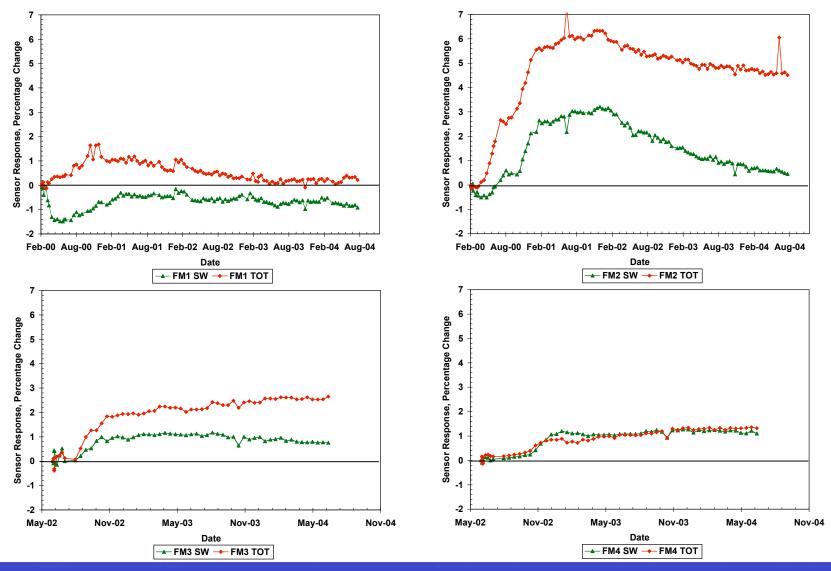








CERES Solar Calibration Results

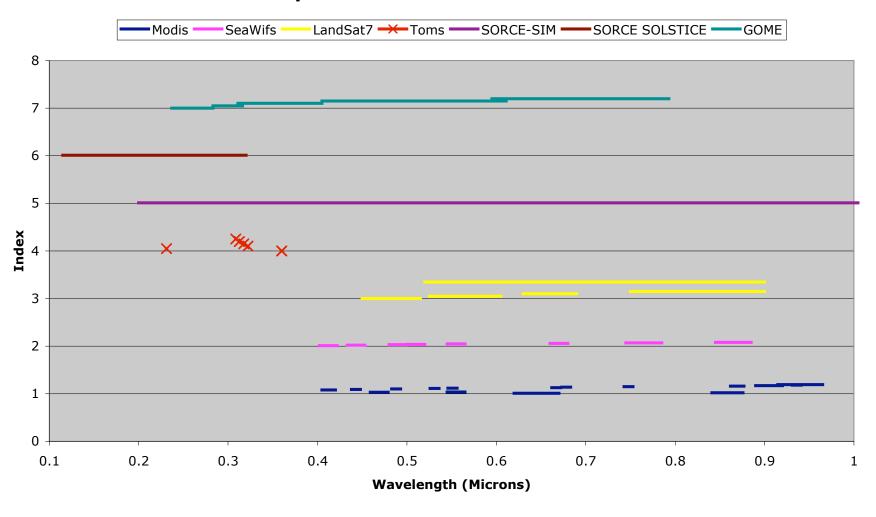






LEO Missions Subject to Spectral Darkening

Bandpasses of Selected Instruments

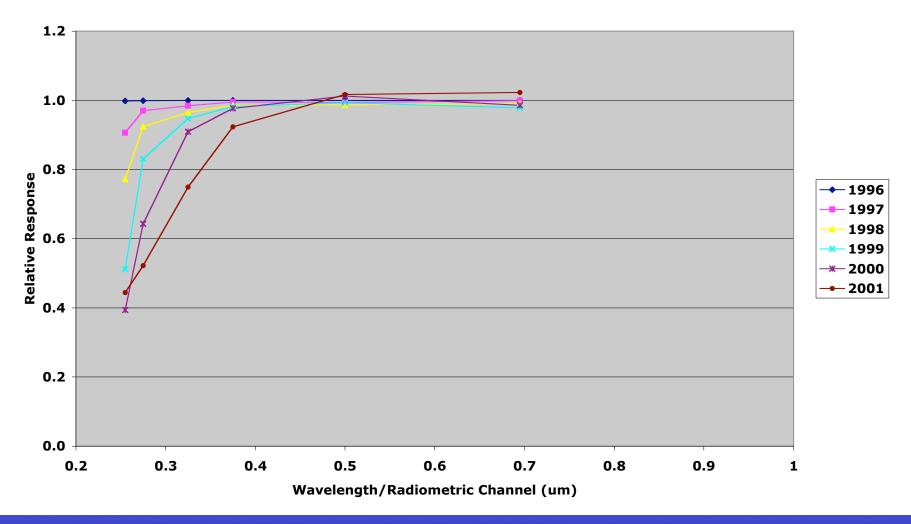






Spectral Darkening on Similar Missions

Global Ozone Monitoring Experiment (GOME) Spectral Darkening



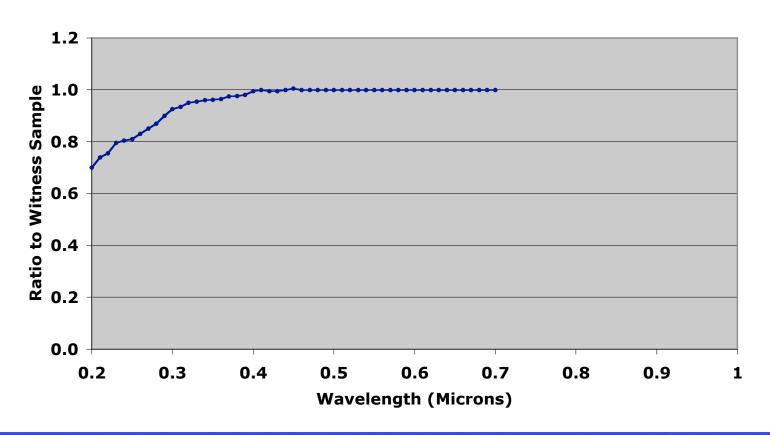




Spectral Darkening on Similar Missions

Spare NIMBUS ERB Radiometer flown on LDEF

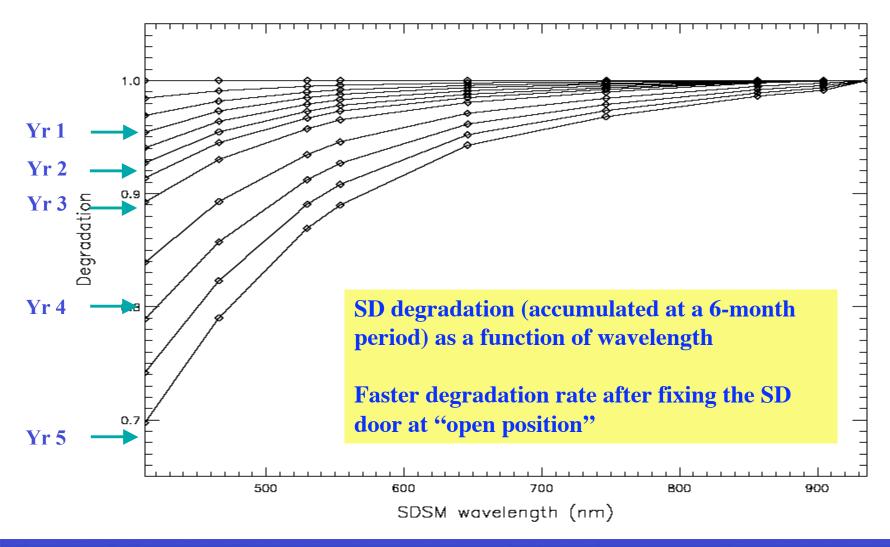
Measured Transmittance Change of Suprasil W 4mm sample, 5 yr. exposure







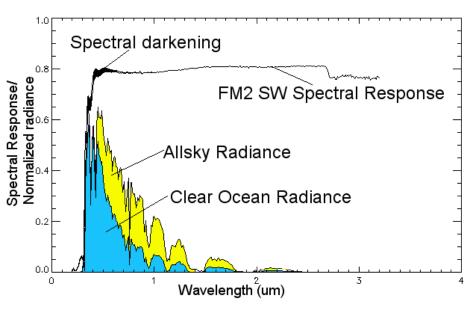
Modis Solar Diffuser Spectral Degradation

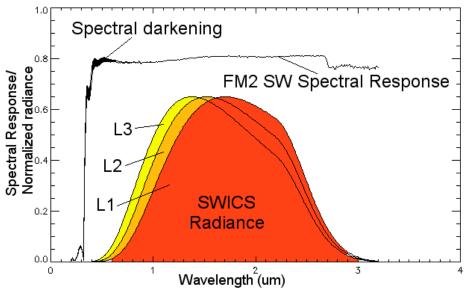






Spectral Degradation: SWICS vs. Earth Spectra





 f_{allsky}^{sw} change $\approx -2\%$

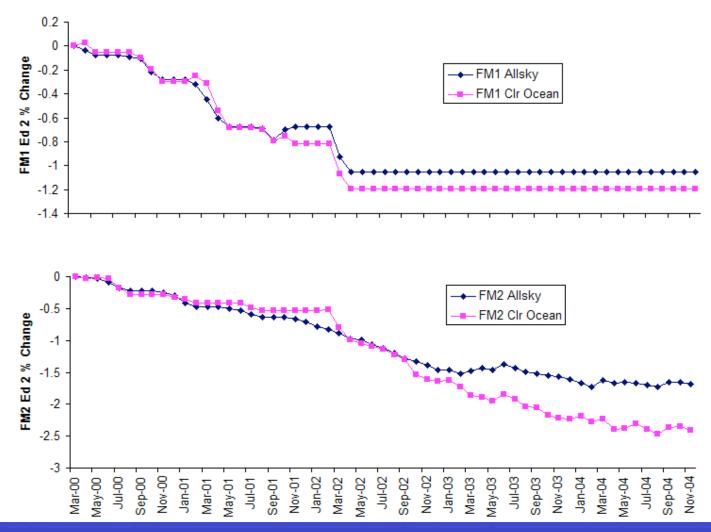
 f_{swics}^{sw} change $\approx -0.1\%$





Edition2_Rev1 SW Scaling Factors

Use Xtrack instrument as calibration reference to derive Rev1 Scaling Factors







Edition2 Rev1 SW Scaling Factors

A table of Rev1 adjustment factors is issued via the quality summary, authors then use the description "Edition2_Rev1"



CERES Terra Revision Table

2000 | 2001 | 2002 | 2003 | 2004 | Tab-delimited file

Feb 1.011 1.015 1.012 1.018 Mar 1.011 1.015 1.012 1.019 Apr 1.011 1.014 1.012 1.019				All Sky	Clear Ocean		
Feb 1.011 1.015 1.012 1.018 Mar 1.011 1.015 1.012 1.019 Apr 1.011 1.014 1.012 1.019	Year	Month	FM1	FM2	FM1	FM2	
Mar 1.011 1.015 1.012 1.019 Apr 1.011 1.014 1.012 1.019	2003	Jan	1.011	1.015	1.012	1.017	
Apr 1.011 1.014 1.012 1.019		Feb	1.011	1.015	1.012	1.018	
		Mar	1.011	1.015	1.012	1.019	
		Apr	1.011	1.014	1.012	1.019	
		May	1.011	1.015	1.012	1.020	
						\ \	

SW fluxes by these values

Multiply your May 03 Terra allsky Multiply your May 03 Terra clear ocean SW fluxes by these values





Terra ES-8 Edition2_Rev1 Summary

•Cal/Val Protocol demonstrates radiometric stability of the data products through 11/2004 of....

	Edition1		Edit	ion2	Edition2_Rev1	
	FM1	FM2	FM1	FM2	FM1	FM2
LWday	.3	.6	.125	.125	.125	.125
LWnight	.1	.125	<.1	<.1	<.1	<.1
sw	.2	.4	.2	.3	.1	.1
WN	<.1	<.1	<.1	<.1	<.1	<.1

Note: Values apply to all-sky global averages

Units are in %/yr





Terra ES-8 Edition2_Rev1 Summary

•Cal/Val Protocol demonstrates radiometric stability of the data products through 11/2004 of....

	Edition1		Edit	ion2	Edition2_Rev1	
	FM1	FM2	FM1	FM2	FM1	FM2
LWday	.3	.6	.125	.125	.125	.125
LWnight	.1	.125	<.1	<.1	<.1	<.1
SW	.2	.4	.2	.3	.1	.1
WN	<.1	<.1	<.1	<.1	<.1	<.1

Note: Values apply to all-sky global averages

Units are in %/yr





CERES Gain/Spectral Change Summary

Cumulative changes incorporated in Edition 2 BDS and ERBE-Like products

			Channe Gain (%)	I	Channel Spectral Response (%)				TOA Flux All-Sky (W/m^2)			
Spacecraft		ТОТ	WN	SW	SW	WN	SW/ TOT	LW/ TOT	SW	LWDay	LWNight	
Terra	FM1	.35	-	-	-	-	.45	-	-	1.1	.65	
	FM2	.40	-	-	.40	-	2.6	-	.4	3.6	1.0	
Aqua	FM3	.45	-	.50	-	-	1.4	-	.5	1.6	1.1	
	FM4	.30	-	.80	-	-	1.2	-	.7	1.6	.75	

- 1. All values are maximum on-orbit changes
- 2. Ground to flight shifts are accounted for separately
- 3. SW/TOT < 3.0 microns, LW/TOT > 3.0 microns
- 4. TOA Flux values are all-sky global averages





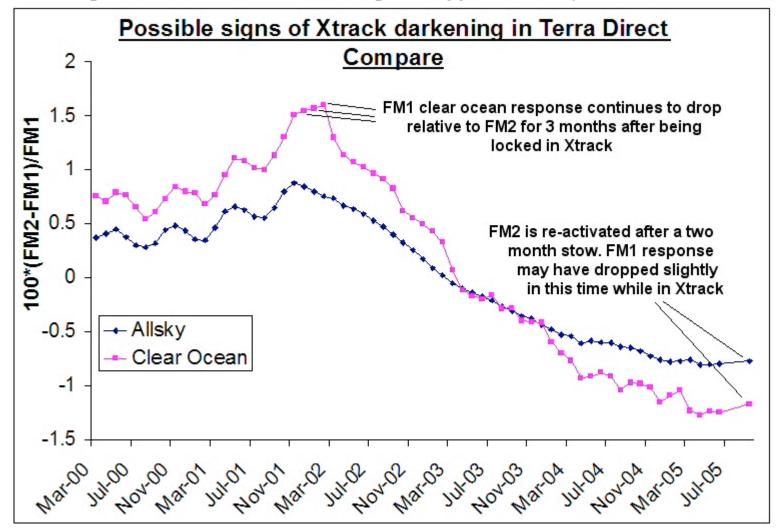
Derivation of Edition 3 SW Spectral Response

CERES Instrument Working Group CERES STM, 1 Nov 05, Hampton, VA





Rev1 adjustment removes the drift assuming the Xtrack instrument is stable. Some unexplained features in direct compare suggest this may not be the case:





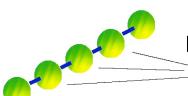


In order to fully compensate for darkening effects in the Edition 3 spectral response the following activities are underway:

- Further stows are planned to directly measure any Xtrack darkening
- New reduced noise full swath direct compare data using both instruments in Xtrack is under production
- A physical model of the contaminant deposition and subsequent darkening is being developed and tuned to internal calibration and direct compare data. This will incorporate results from further stows and noise reduced direct compare





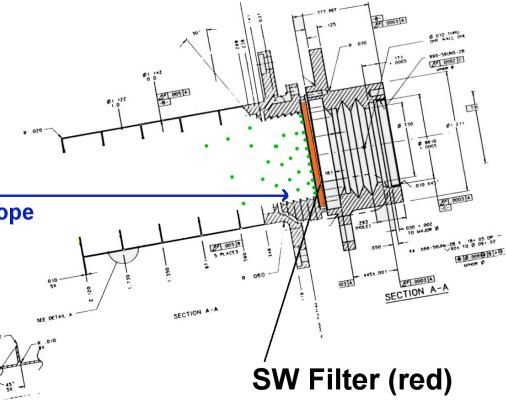


Polymerized contaminant molecules in absorbing chain, category 'B'

Un-polymerized 'solo' contaminant molecule category 'A' (non-absorbing)

Space bound particles enter telescope when in RAM direction, freeing contaminant molecules to be fixed to filter

by Earth UV







Photons tend to interact with targets of similar size to their wavelength, hence contaminant absorption can be modeled as:

 $a(\lambda) = H \times \text{(number of of molecules in a polymer chain of length } \lambda)$

Might expect population density of polymer chain length to fall off exponentially:

$$a(\lambda) = H \times e^{-\alpha \lambda}$$

 $T(\lambda) = 1 - a(\lambda)$

$$T(\lambda) = 1 - a(\lambda)$$

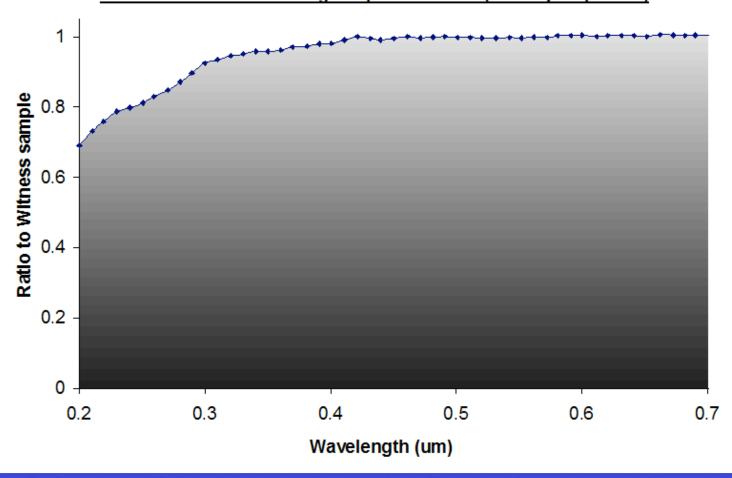
$$T(\lambda) = 1 - H.e^{-\alpha\lambda}$$





This fits the basic shape of spectral darkening seen on LDEF, GOME and MODIS: $T(\lambda) = 1 - H.e^{-\alpha\lambda}$

LDEF Transmittance Change Suprasil W 4mm (CH5 6 yr exposure)

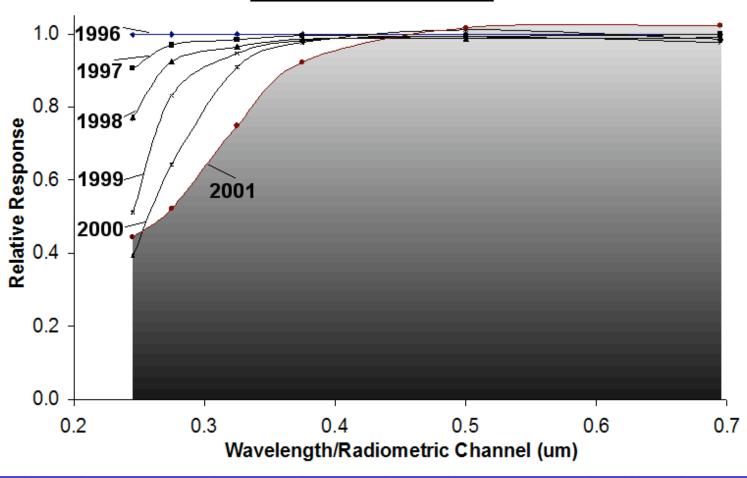






This fits the basic shape of spectral darkening seen on LDEF, GOME and MODIS: $T(\lambda) = 1 - H.e^{-\alpha\lambda}$

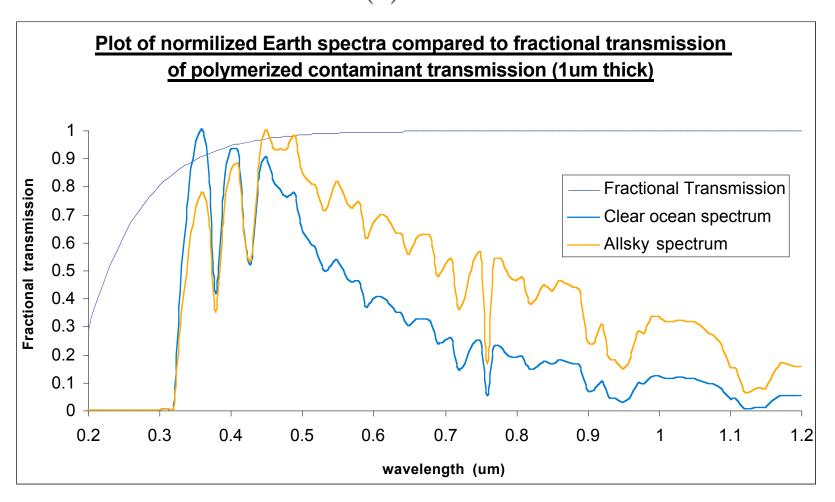
GOME Spectral Darkening







So 1 um thickness of polymerized contaminant might have a transmission given by: $T(\lambda) = 1 - 0.71.e^{-12.95\lambda}$







Experience from LDEF suggests that continued exposure of contaminants to UV photons increases the polymerization, making the remaining category 'A' solo molecules form more absorbing long chains (becoming category 'B')

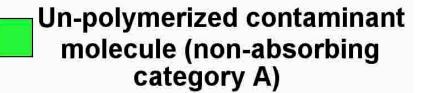
$$\frac{dB}{dt} = 0.008 [N - B] + 0.4 \frac{dN}{dt}$$

If N is the total number of contaminant molecules (A+B), this equation represents a situation where 40% of molecules arriving per day of RAPs are already polymerized (category B). Every day earth UV also converts 0.8% of the remaining category A molecules to category B









RAPS

Contaminant Thickness





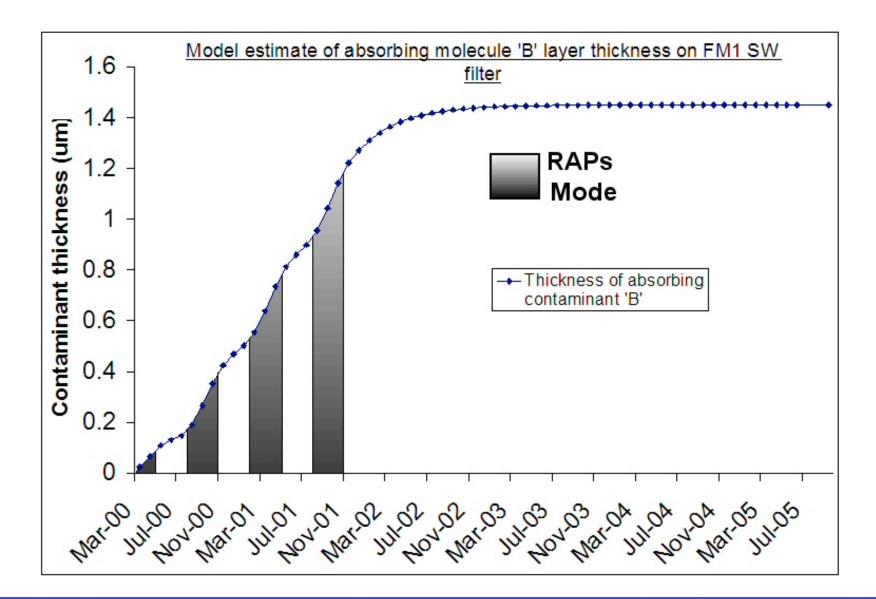
Read in the modes for each instrument and use them to generate forcing (dN/dt) for arrival of contaminant at each instrument filter surface



DATE	FM1_INSTRUMENT		FM2_INSTF	RUMENT	r					
MARCH 2	000									
03/01/00	RAPS	1	XTRACK	0						
03/02/00	XTRACK	0	XTRACK	0						
03/03/00	XTRACK	0	RAPS	1						
03/04/00	RAPS	1	XTRACK	0						
03/05/00	XTRACK	0	XTRACK	0						
03/06/00	XTRACK	0	RAPS	1						
03/07/00	RAPS	1	XTRACK	0						
	XTRACK	0	XTRACK	0						
03/09/00	XTRACK	0	RAPS	1						
03/10/00	RAPS	1	XTRACK	0	(SunRise	SolCal:	FM1-00:43	UTC,	FM2-01:28	UTC)
03/11/00	XTRACK	0	XTRACK	0	(SunRise	SolCal:	FM1-01:28	UTC,	FM2-01:26	UTC)
03/12/00	XTRACK	0	RAPS	1						
03/13/00	XTRACK	0	XTRACK	0						
03/14/00	XTRACK	0	XTRACK	0						
03/15/00	XTRACK	0	XTRACK	0						
03/16/00	RAPS	1	XTRACK	0	(SunRise	SolCal:	FM1-00:06	UTC,	FM2-00:08	UTC)
03/17/00	XTRACK	0	XTRACK	0						
03/18/00	XTRACK	0	RAPS	1						
03/19/00	RAPS	1	XTRACK	0						
03/20/00	XTRACK	0	XTRACK	0						
03/21/00	XTRACK	0	RAPS	1						
03/22/00	RAPS	1	XTRACK	0						
03/23/00	XTRACK	0	XTRACK	0						
03/24/00	XTRACK	0	RAPS	1						
03/25/00	RAPS	1	XTRACK	0						
03/26/00	XTRACK	0	XTRACK	0						
03/27/00	XTRACK	0	RAPS	1						
03/28/00	RAPS	1	XTRACK	0						
03/29/00	XTRACK	0	XTRACK	0						
03/30/00	XTRACK	0	RAPS	1						
03/31/00	RAPS	1	XTRACK	0						
APRIL 2	000									
04/01/00	RAPS	1	XTRACK	0						
04/02/00	RAPS	1	XTRACK	ō						
04/03/00	RAPS	1	XTRACK	ō						
04/04/00	RAPS	1	XTRACK	0						
04/05/00	RAPS	1	XTRACK	0						

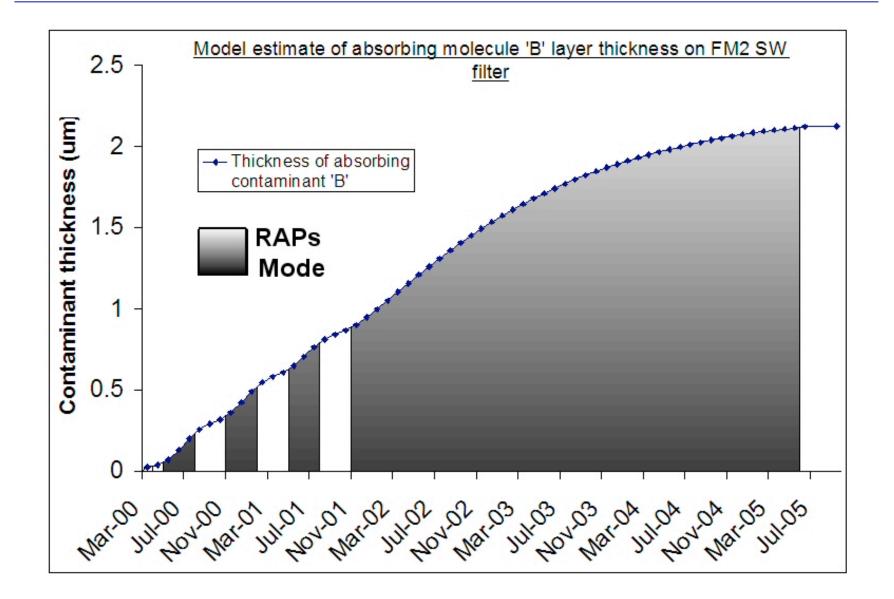






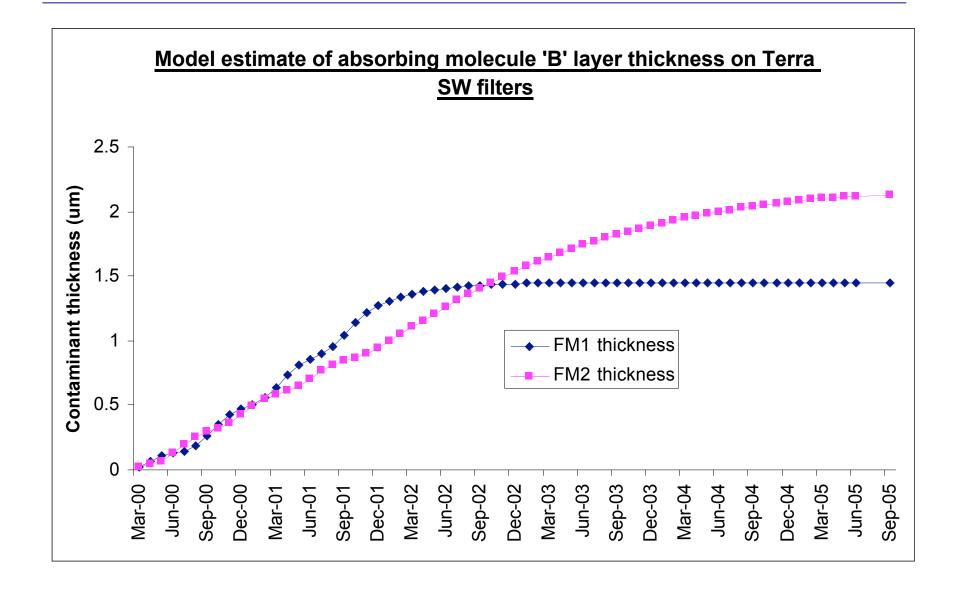








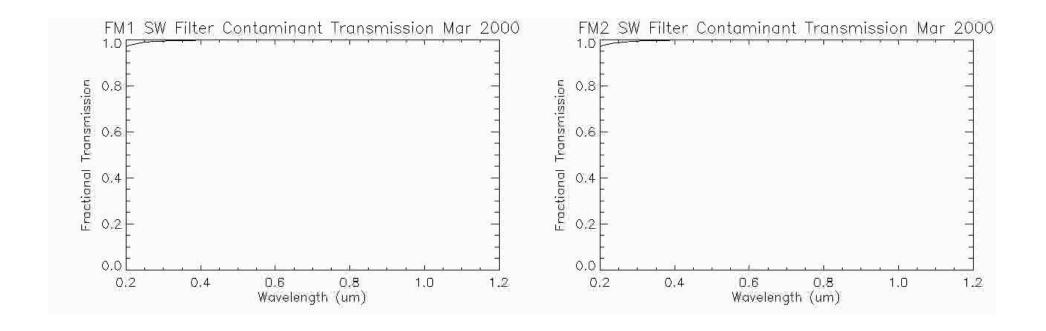








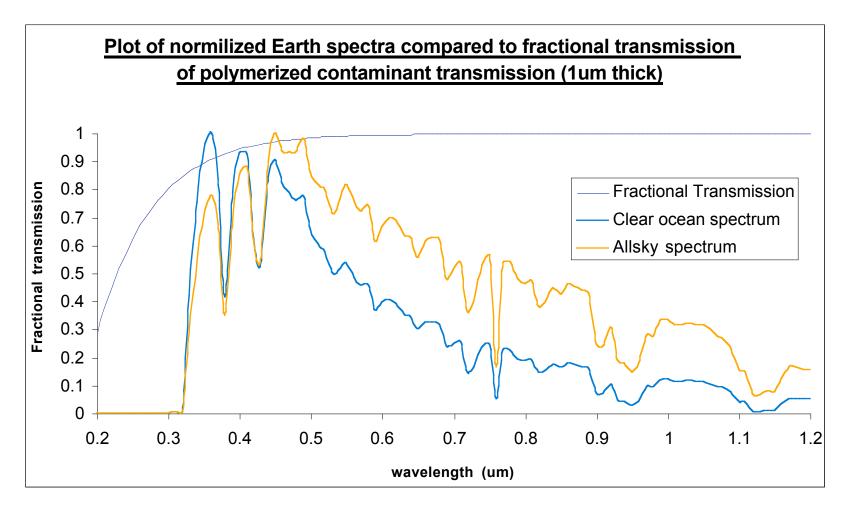
Hence the contaminant transmission can now be found throughout the mission: $T(\lambda) = \left[1 - 0.71.e^{-12.95\lambda}\right]^{thickness}$





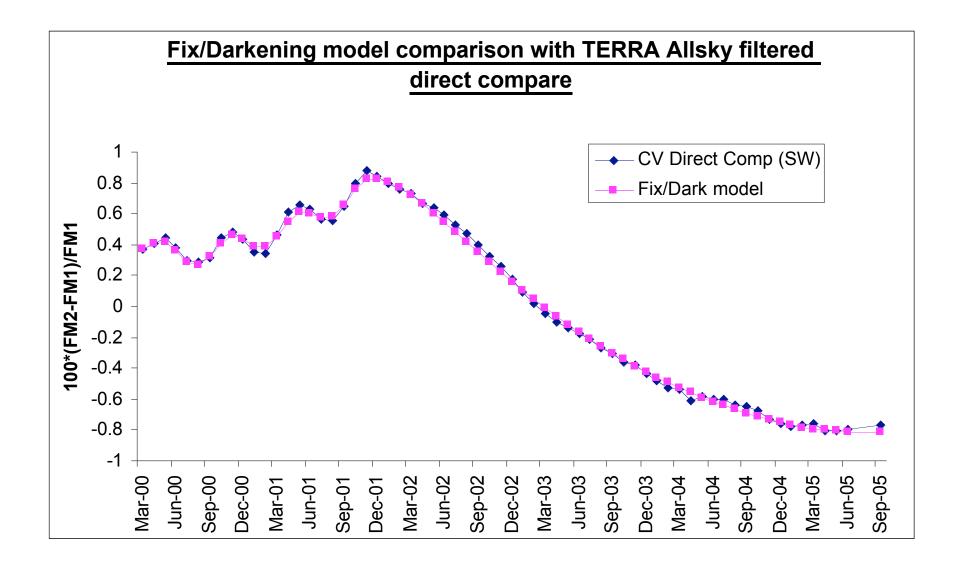


Tune the model so its estimate of changes to filtered radiance matches the internal calibration lamps and direct compare:



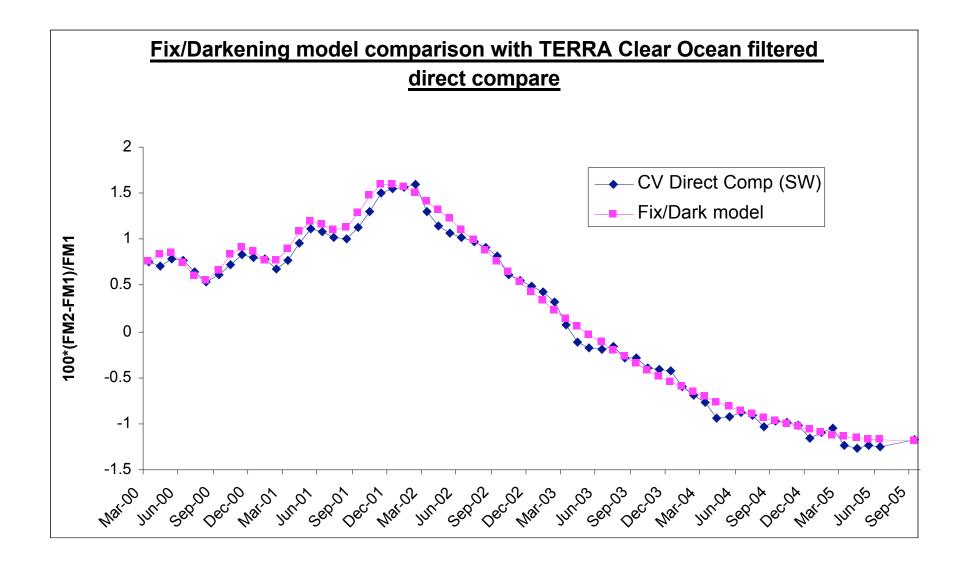








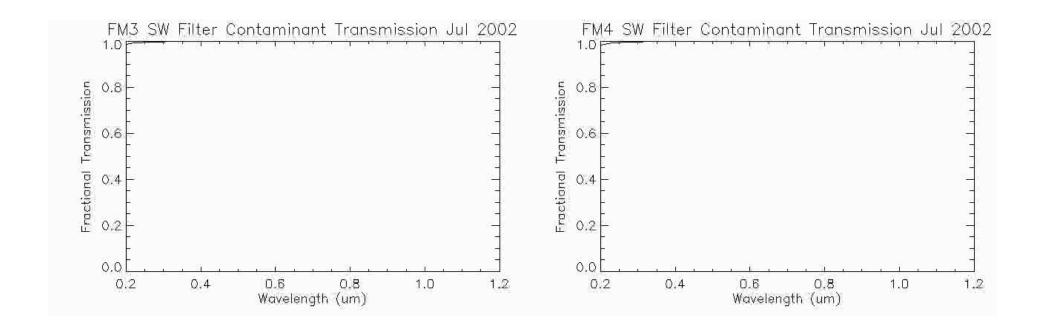






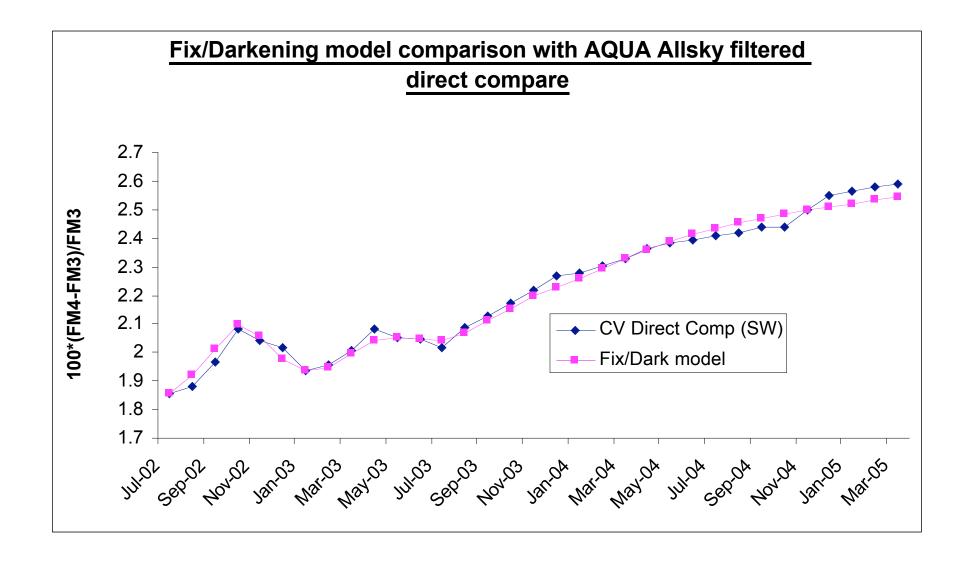


The model can also be tuned to Aqua direct compare and internal calibration results:













Summary and Conclusions

- Rev 1 adjustments remove the majority of the spurious SW trends in Edition 2 data
- Special operations and modeling currently underway to characterize and compensate for RAPs and any Xtrack darkening in Edition 3 SW spectral response
- With SW channel model satisfactorily tuned, existing 3 channel inter-comparison method can be modified to derive any spectral darkening to total channel optics



